Black holes and random matrices

Stephen Shenker

Stanford University

Strings 2017

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- What accounts for the finiteness of the black hole entropy-from the bulk point of view?
- The stakes are high here. Many approaches to understanding the bulk–
 - TFD/Eternal black hole
 - Ryu-Takayanagi
 - Geometry from entanglement
 - Tensor networks
 - ER = EPR
 - Code subspaces
 - ...

suggest that any complete bulk description of quantum gravity must be able to describe these states.

 A simple diagnostic of a discrete spectrum [Maldacena]. Long time behavior of (O(t)O(0)). (O is a bulk (smeared boundary) operator)

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(See also [Dyson-Kleban-Lindesay-Susskind; Barbon-Rabinovici])

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• To focus on the oscillating phases remove the matrix elements. Use a related diagnostic: [Papadodimas-Raju]

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• The "spectral form factor"

Properties of $Z(t)Z^*(t)$

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$$Z(\beta, 0)Z^*(\beta, 0) = Z(\beta)^2 \ (= L^2 = e^{2S} \text{ for } \beta = 0)$$

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ightarrow e^{S}$, an exponential change. How does this occur?

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- G(t, t'), Σ(t, t') description has aspects reminiscent of a bulk description:
 - O(N) singlets
 - nonlocal
 - Nonperturbatively well defined (two replicas)

$$\langle Z(t)Z^{*}(t) \rangle = \int dG_{ab}d\Sigma_{ab} \exp(-N \ I(G_{ab},\Sigma_{ab}))$$

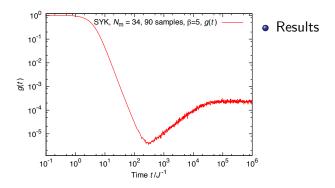
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- Finite dimensional Hilbert space, $D = L = 2^{N/2}$, amenable to numerics
- Guidance about what to look for

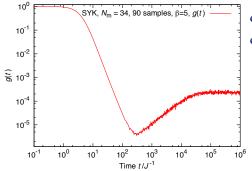
[Jordan Cotler, Guy Gur-Ari, Masanori Hanada, Joe Polchinski, Phil Saad, Stephen Shenker, Douglas Stanford, Alex Streicher, Masaki Tezuka] ([CGHPSSSST])

See also [Garcia-Garcia-Verbaarschot]

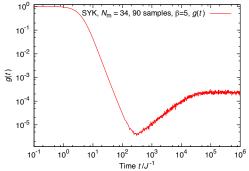


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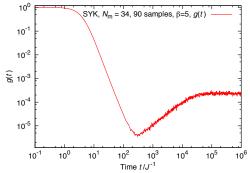
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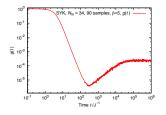
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- The Slope ↔ Semiclassical quantum gravity



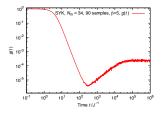
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- The Dip \leftrightarrow crossover time



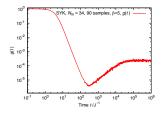
Slope is determined by semiclassical quantum gravity-nonuniversal

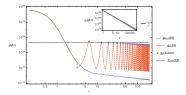


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In SYK slope $\sim 1/t^3$. One loop exact Schwarzian result: $\rho(E) \sim e^{S_0}(E - E_0)^{1/2}$ ([Bagrets-Altland-Kamenev; CGBPSSSST; Stanford-Witten])



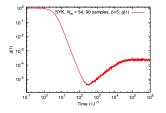


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In BTZ summing over modular transforms of blocks gives oscillating slope with power law envelope: nonperturbatively small oscillations in the density of states.

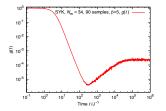
[Dyer-Gur-Ari]



The Ramp and Plateau are signatures of Random Matrix Statistics, believed to be universal in quantum chaotic systems



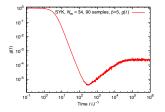
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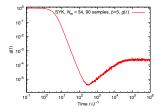
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[Dyson; Gaudin; Mehta]





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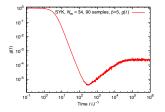
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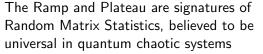
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The decrease before the plateau is due to anticorrelation of levels







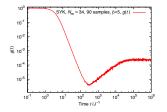
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Some evidence for this in melonic models [Witten; Gurau; Carrozza-Tanasa;

Klebanov-Tarnopolsky; Krishnan-Kumar=Sanyal]

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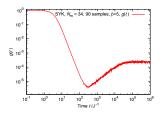
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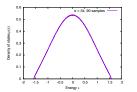
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- q = 2 SYK in progress [Saad, SS]

[Gharibyan-Hanada-SS-Tezuka, in progress]

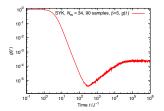


At what time does the ramp begin?

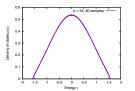


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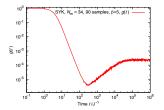
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At what time does the ramp begin? The dip is just a crossover: edge versus bulk dynamics

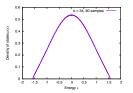


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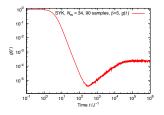


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The Thouless time [Garcia-Garcia-Verbaarschot]



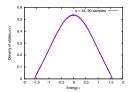
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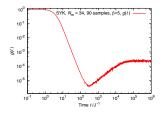
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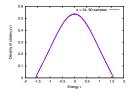
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Single particle hopping, *n* sites: diffusion time, $t_{th} \sim n^2$



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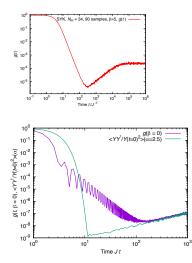
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Follow the ramp below the slope: use Gaussian filter [Stanford]

$$Y(\alpha,t)Y^*(\alpha,t) = \sum_{m,n} e^{-\alpha(E_n^2 + E_m^2)} e^{+i(E_m - E_n)t}$$

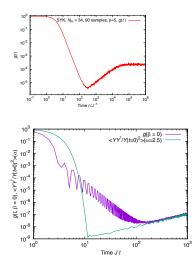
 YY^*



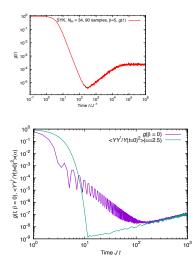
Dip time $t_d \sim 200$, N = 34

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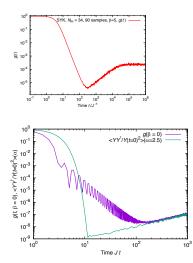


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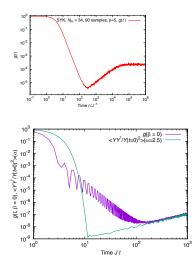
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Onset of ramp $t_r \lesssim 10$, N = 34

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An upper bound. Very little variation in N for $N \leq 34$

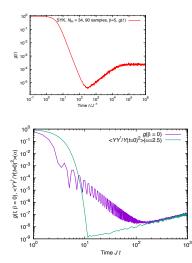


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log N? scrambling?



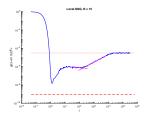
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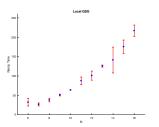
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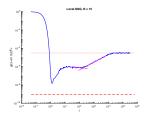
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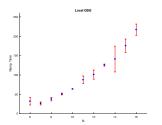
Maybe; no.



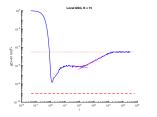


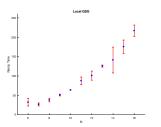
n geometrically local qubits





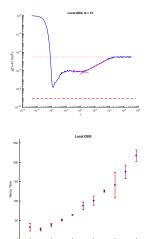
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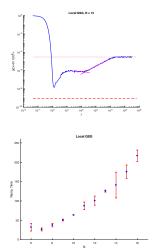
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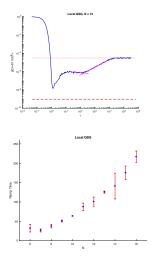
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slope $\sim \exp(-Nt^2)$, rapid decay

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- Can analyze dynamics including scrambling analytically [Oliveira-Dahlsten-Plenio; Lashkari-Stanford-Hastings-Osborne-Hayden; Harrow-Low; ...]

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- Time to randomize last qubit ~ *n*, scrambling time [Nahum-Vijay-Haah; Keyserlingk-Rakovszky-Pollmann-Sondhi]

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- Correlation functions of very complicated operators [Roberts-Yoshida; Cotler-Hunter-Jones-Liu-Yoshida]

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- So these phenomena would appear in small black holes as well, although as an exponentially subleading effect
- We need to know what they mean in quantum gravity!