

What has been learned from BICEP2?

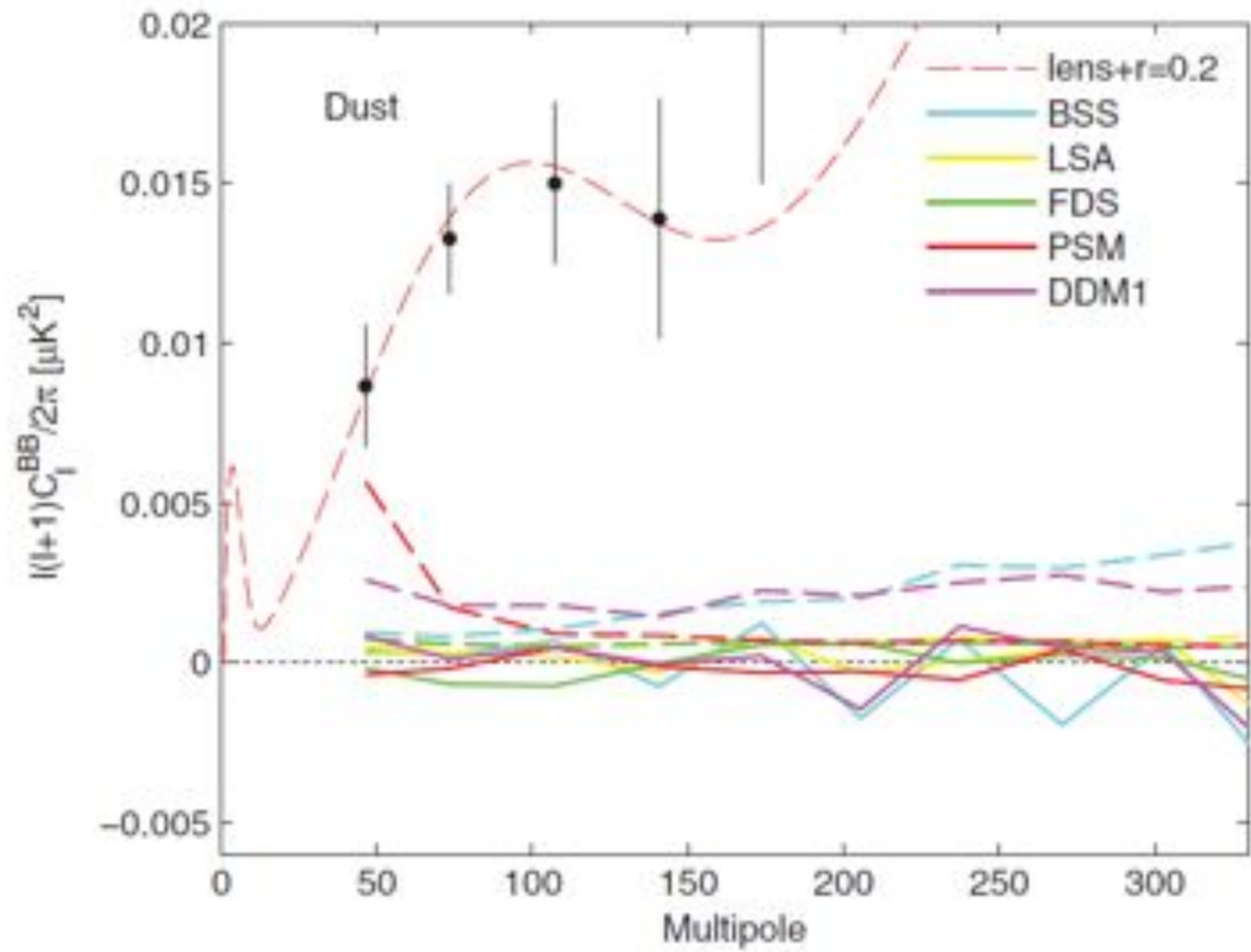
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Step 1. Evaluate the null hypothesis

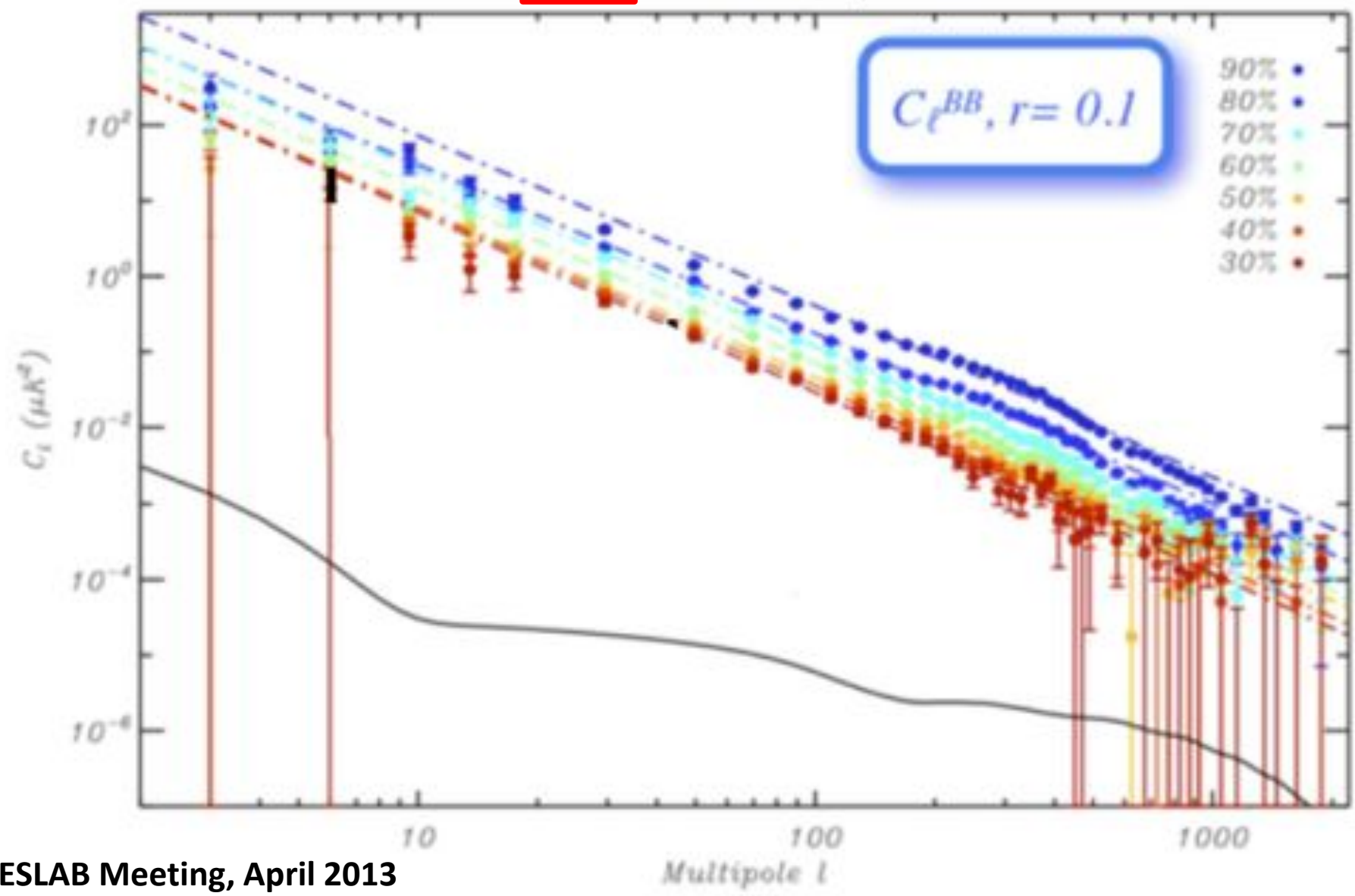
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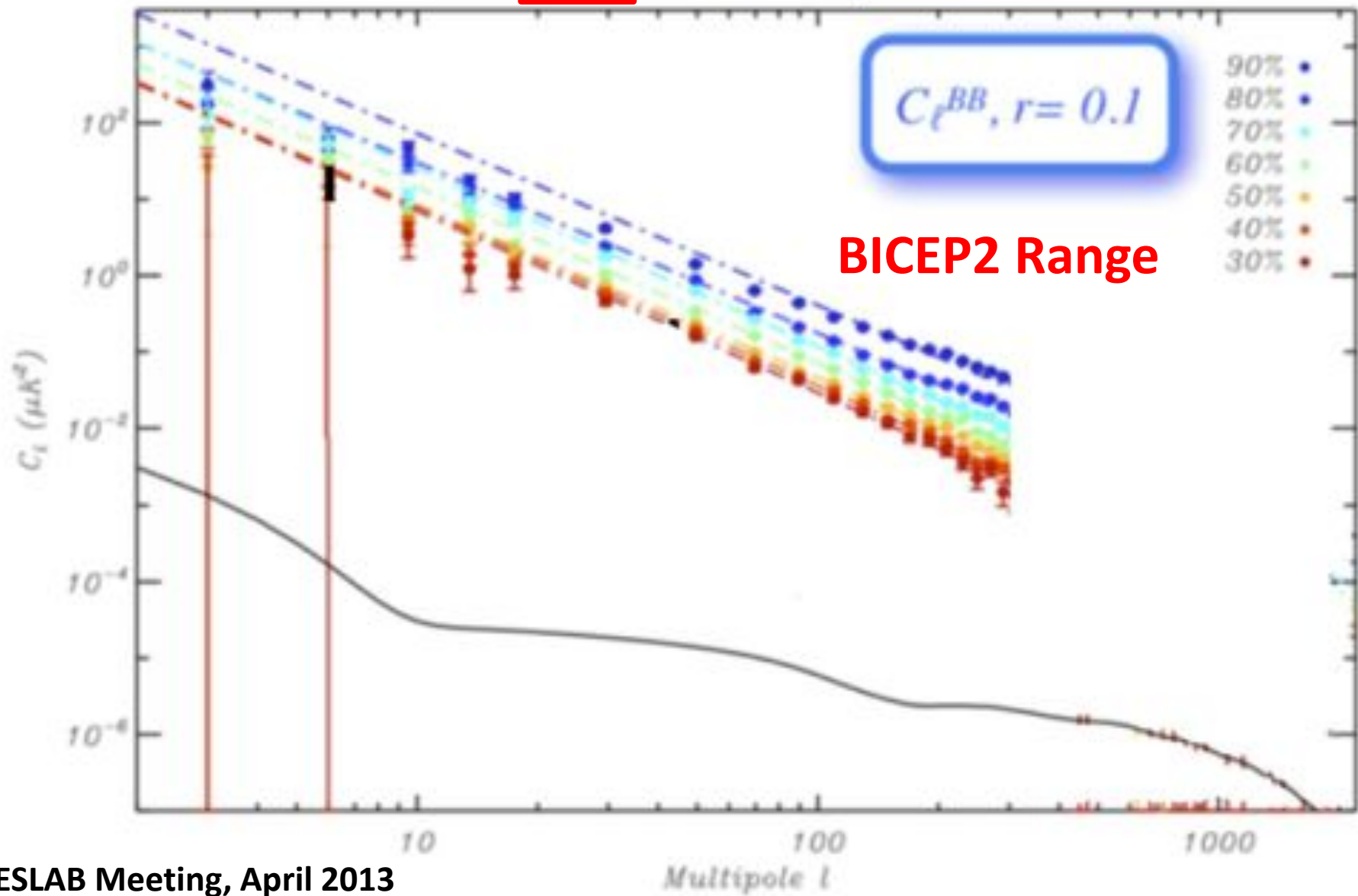
The signal detected by BICEP2 is due to
lensing + foreground + instrumental noise



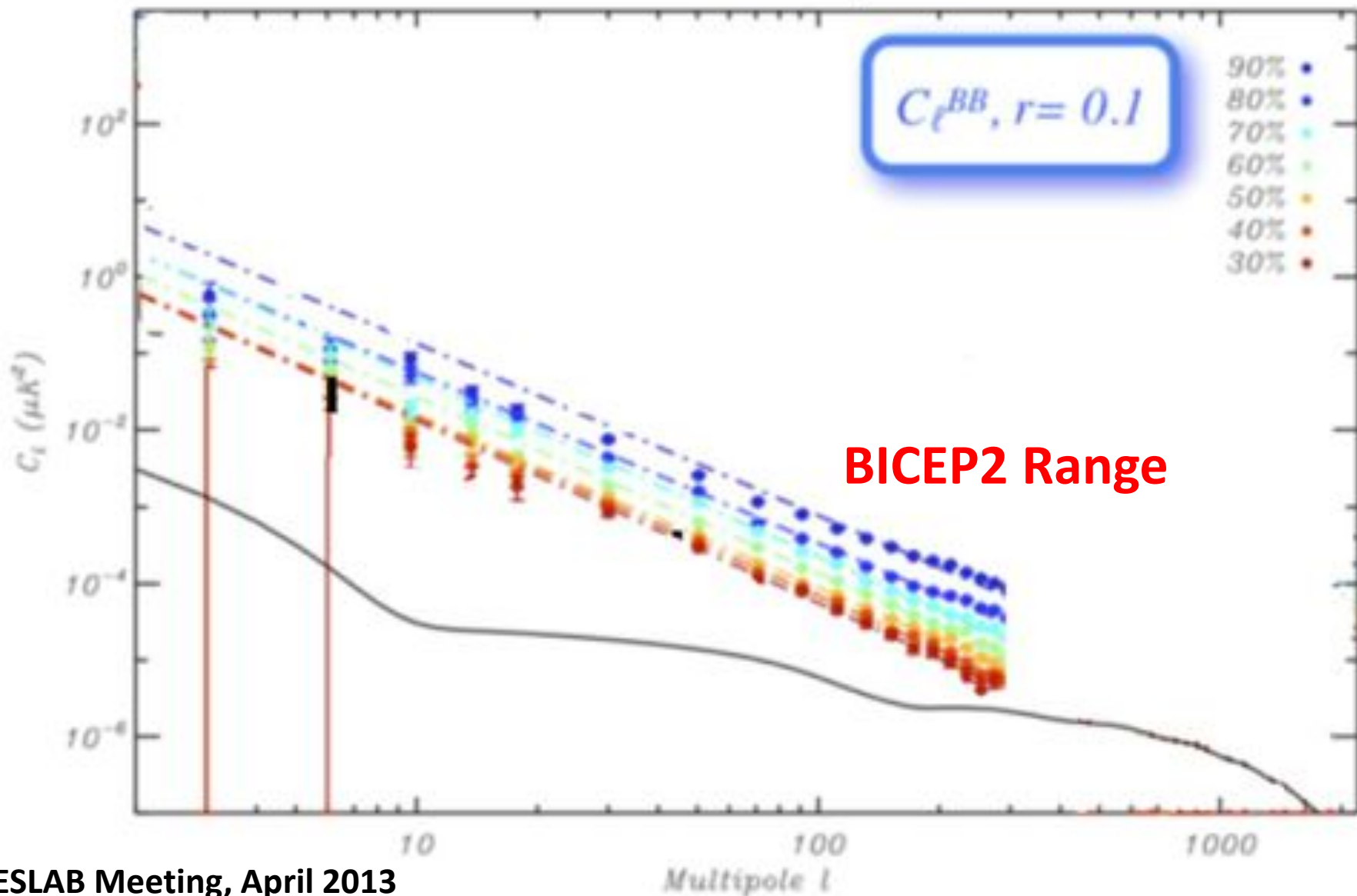
353 GHz BB spectrum

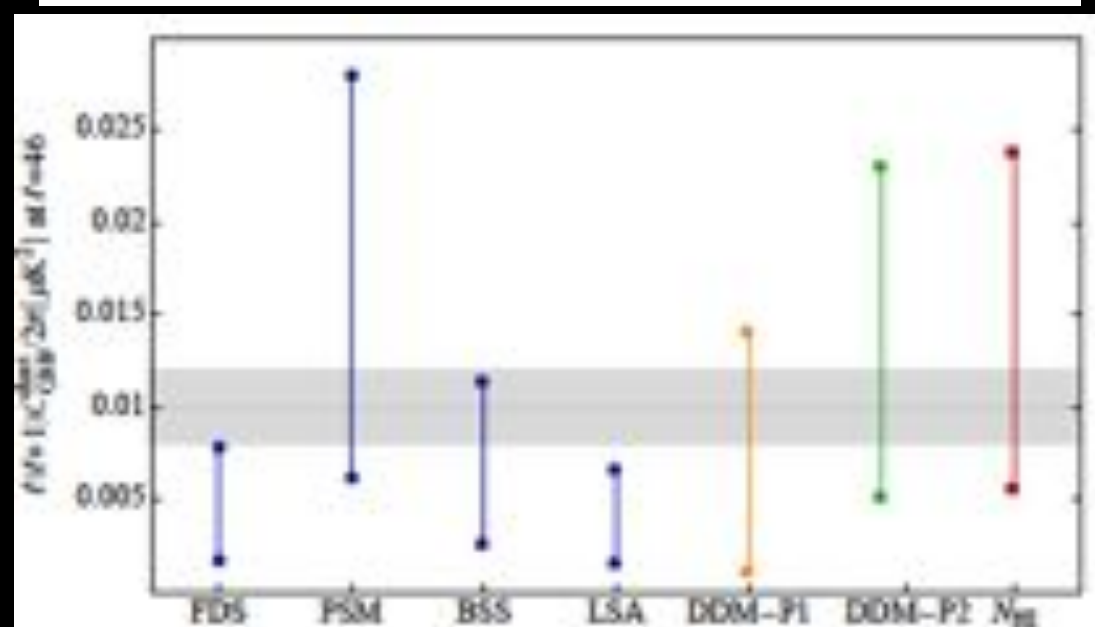
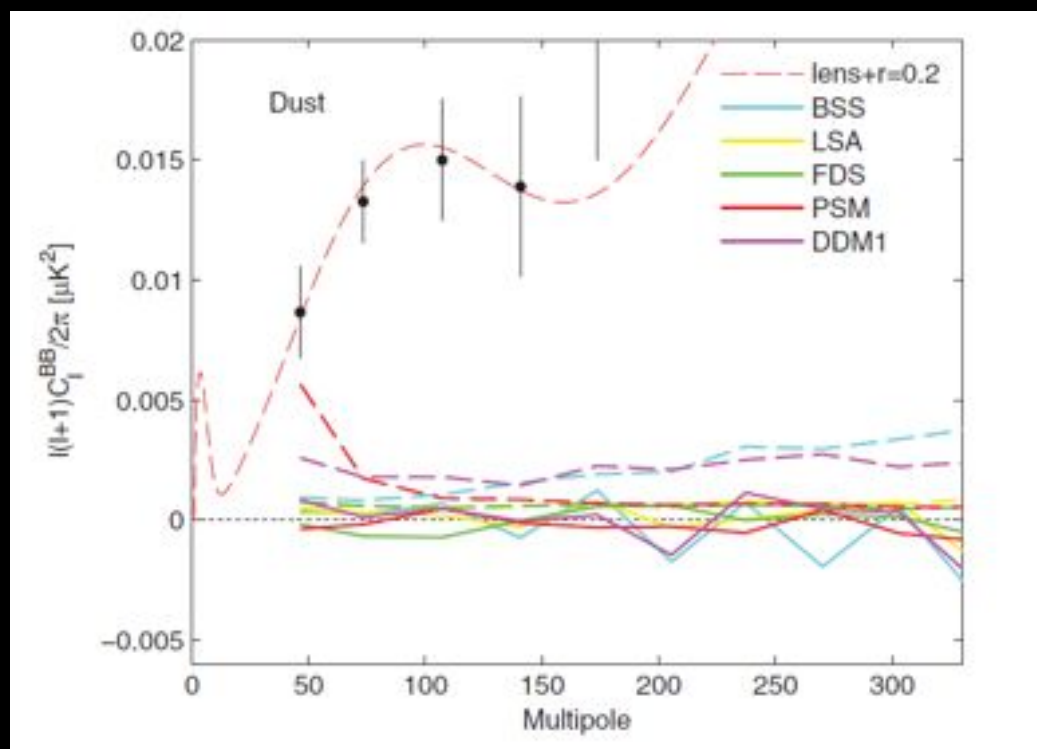


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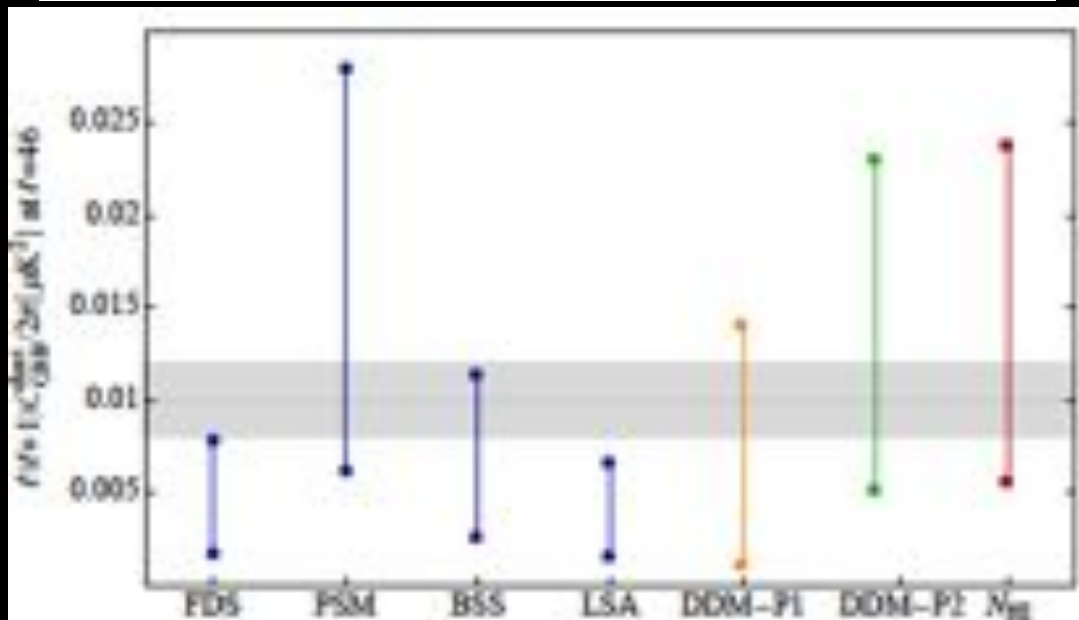
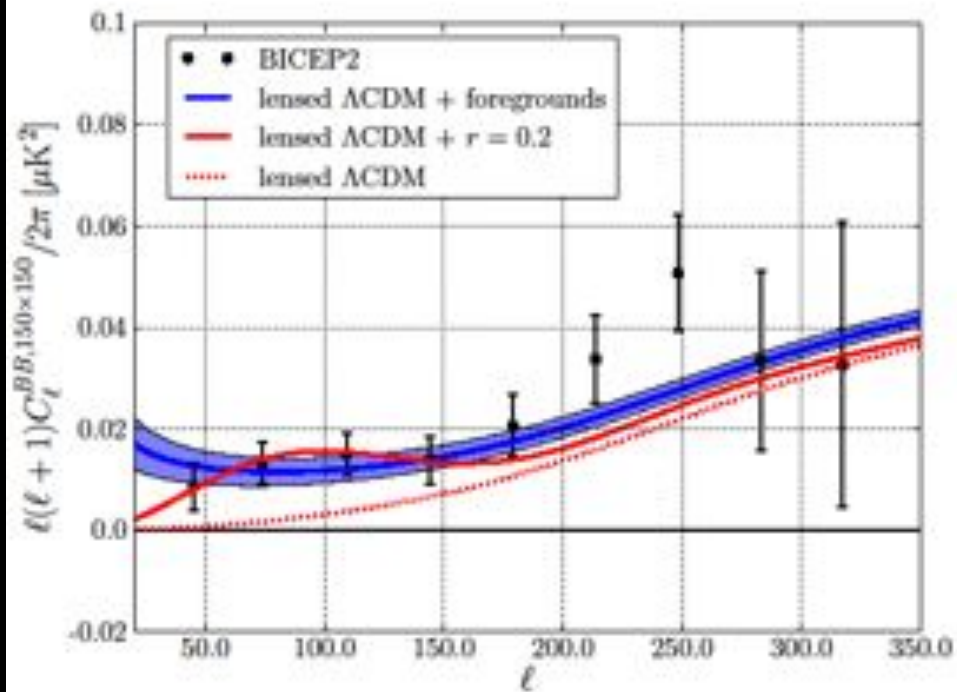


143 GHz BB spectrum





Flauger, Hill and Spergel, arXiv:1405.7351



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Conclusion:

Based on the best available data at the time (and now),
the null hypothesis could not be rejected

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The detected signal is due to
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Step 2. If null hypothesis is not rejected, do not
proceed to step 3.

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Do not consider the ALTERNATIVE hypothesis
(that the signal is due to gravitational waves)

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$$r = 0.2^{+0.07}_{-0.05} \quad \text{ruling out } r=0 \text{ at } 7\sigma \quad ??$$



What has been learned from BICEP2?

The inflationary paradigm is so flexible
that *no test or combination of tests*
can disprove it

Note: It makes no difference if individual versions of a paradigm are testable.

If the paradigm includes a spectrum of versions that produce every conceivable outcome, then the paradigm is untestable.

Feynman (Cornell, 1964): The Scientific Method
(see also Feynman's lecture entitled "Cargo Cult Science")

*The vagueness of the inflationary paradigm
is not disputed by proponents
– rather, it is embraced*

Harvard CMB Symposium (2014)

Q: Is inflation falsifiable?

Alan Guth: “I think that is kind of a silly question.

... I think inflation is too flexible of an idea
for that to make sense.”

Andrei Linde, March 17 press release, Stanford U

“These results are a smoking gun for inflation, because alternative theories do not predict such a signal,” he said. “This is something I have been hoping to see for 30 years.”

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- *outcome sensitive to initial conditions*
- *outcome sensitive to parameters*
- *outcome varies across the multiverse*

Must Rethink

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solve the initial conditions
& multiverse problems of inflation

find an alternative paradigm

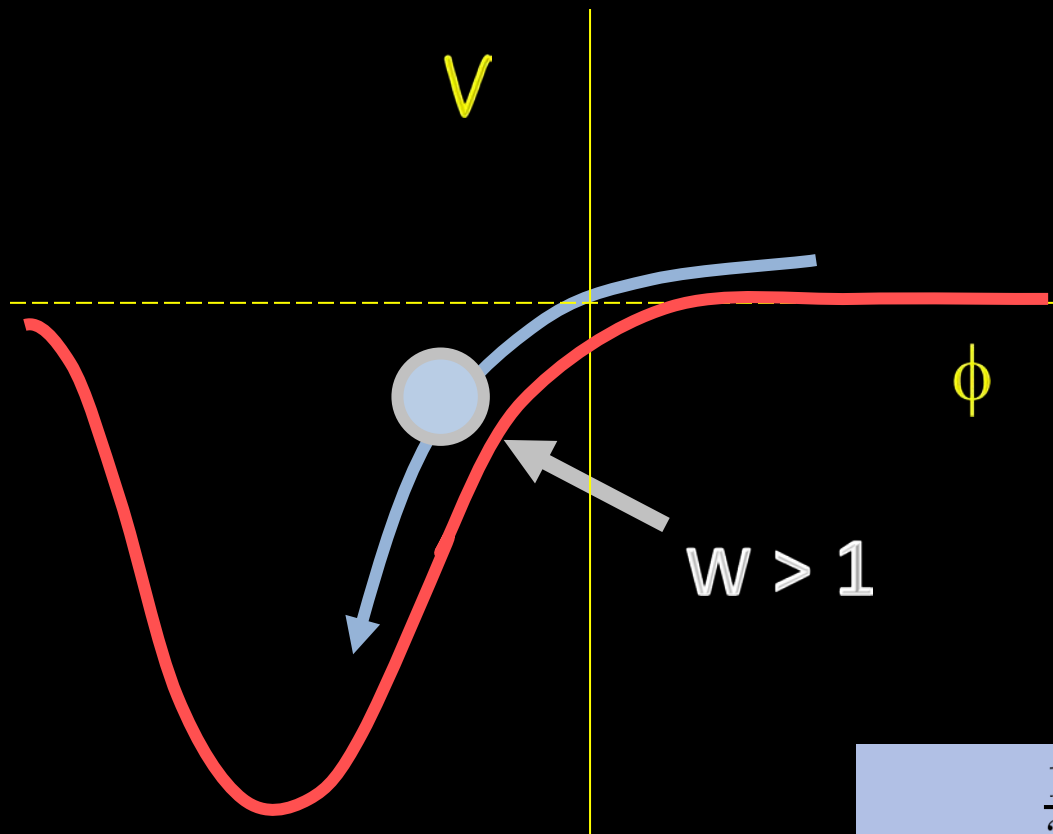
why consider a bounce?

simple solution to flatness & homogeneity problems

$$H^2 = \frac{8\pi G}{3} \frac{\rho_m^0}{a^3} + \frac{8\pi G}{3} \frac{\rho_r^0}{a^4} + \frac{\sigma^2}{a^6} + \dots - \frac{k}{a^2}$$
$$+ \frac{8\pi G}{3} \frac{\rho_\phi^0}{a^{3(1+w)}} \leftarrow w = p/\rho \geq 1$$

“EKPYROTIC”

ultra-high pressure, ultra-slow contraction



$$w = \frac{\frac{1}{2}\dot{\phi}^2 - V(\phi)}{\frac{1}{2}\dot{\phi}^2 + V(\phi)}$$

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evades the multiverse/unpredictability problem

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nearly scale-invariant density fluctuations
but no observable tensors ($r \approx 0$)

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cyclic

classically geodesically complete

current vacuum metastable or unstable

What happened at the bang?



What happened at the bounce?

a proposal

Note that, near the bounce, the effective action simplifies:

$$S = \int d^4x \sqrt{-g} \left[\frac{1}{2\kappa^2} R(g) - \frac{1}{2} (\partial\sigma)^2 + \text{radiation} \right]$$

... so can reformulate in Weyl-invariant form:

$$S = \int d^4x \sqrt{-g} \left[\frac{(\partial\phi)^2 - (\partial s)^2}{2} + \frac{\phi^2 - s^2}{12} R + \text{radiation} \right]$$

... now classical solutions for ϕ and s can be found that continue through the bounce, from big crunch to big bang, and cyclic cosmology can be made geodesically complete!

*Weyl invariance can be extended
to all currently known physics*

$$\mathcal{L}(x) = \begin{aligned} & \frac{1}{12} (\phi^2 - 2H^\dagger H) R(g) \\ & + g^{\mu\nu} \left(\frac{1}{2} \partial_\mu \phi \partial_\nu \phi - D_\mu H^\dagger D_\nu H \right) \\ & - \left(b\phi^4 + \frac{\lambda}{4} (H^\dagger H - \xi^2 \phi^2)^2 \right) \\ & + L_{\text{SM}} \left(\begin{array}{l} \text{quarks, leptons, gauge bosons,} \\ \text{Yukawa couplings to } H \end{array} \right) \end{aligned}$$

$$\begin{aligned} g_{\mu\nu} &\rightarrow \Omega^{-2} g_{\mu\nu}, \quad \phi \rightarrow \Omega \phi, \quad H \rightarrow \Omega H, \\ \psi_{q,l} &\rightarrow \Omega^{3/2} \psi_{q,l}, \quad A_\mu^{\gamma,W,Z,g} \rightarrow \Omega^0 A_\mu^{\gamma,W,Z,g} \end{aligned}$$

$$H(x) = \begin{pmatrix} 0 \\ \frac{1}{\sqrt{2}} |s(x)| \end{pmatrix}$$

can also add dark matter and r-handed ν 's

Bars, PJS, Turok (to appear)

can reformulate string theory with a local scale symmetry in target space without any fundamental lengths such that the fundamental length in string theory – the string tension – emerges from gauge fixing a field.

What happened at the bang?



What happened at the bounce?